

Maximum-Resolution Solar Spectrometer (QMAX[™])

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The Questar Corporation of New Hope, Pennsylvania, is pleased to announce the world's first, Maximum-Resolution Solar Spectrometer (**QMAX**^M) capable of fitting comfortably in the palm of a hand. The Questar **QMAX**^M attaches easily to any telescope that is compatible, according to its manufacturer, with solar observation and that has a 1.25" drawtube. The Questar Maximum-Resolution Solar Spectrometer blocks ultraviolet and infrared radiation while allowing the sun's visible light, the visible solar spectrum, to be viewed in true and gorgeous color.

- Variability
 - Prominence Emission, Radial Height & Structure Hα, Hβ, Hγ, D₃
 - Line reversal over flares emission breaking out abruptly in absorption
 - Doppler shifts over flares velocity-shifted emission to red / blue
 - Monster-sunspot Zeeman splitting Fe I
 - Sunspot line broadening
 - Helium emission from chromosphere, tangential slit, coronal hole indicator
 - Telluric lines: dramatic intensification at Sunset and Sunrise H₂O

Resolution <0.17Angstrom Units (Å) = 0.017nm

- Heisenberg Uncertainty Principle only visible effect of it, Mg Lines
- Molecular bands super-close methyne (CH) energy levels; G-band
- \blacksquare H α so broad that "it looks like a tree trunk"
- Sodium D-Lines so far apart telluric lines between them look like forest
- Line shape sharp vs. broad line wings
- Huge H & K lines

Intimate, beautiful, greatly enlarged views of atomic and molecular lines throughout the solar spectrum at a level of breathtaking detail never before available from such an extremely compact, commercial Spectrometer are made possible by a uniquely Questar insight into the nature of image formation by high-precision optics. US and foreign patents pending.

Orders are currently being taken for June 2002 delivery. http://www.QuestarCorporation.com ©2000 Questar Corporation

Specifications

Maximum Resolution:

Wavelength range:

Ocular:

0.17 Å (Angstrom units), or 0.017 nm (nanometers); see note 1 below

3850 to 6900 Å, or 385 to 690 nm

68 Å, or 6.8 nm at 5500 Å

spectrum formed by the

see note 2 below

see note 3 below

(2Å) increments

95 x relative to the image of the

spectrometer's imaging mirror;

521 x relative to the sun seen

angular, North-South extent appears increased 521 times;

digital readout in 0.2 nanometer

focal plane of QUESTAR'S 8mm Brandon ocular, at 5500 Å.

QUESTAR[®] Brandon 8mm

Width of Field:

Magnification of Spectrum:

Equivalent Magnification of Sun:

Wavelength scale:

Dispersion:

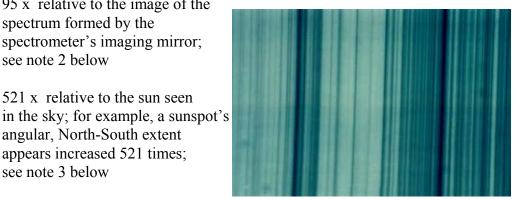
Telescope Pre-filter:

Post-dispersion Filter[.]

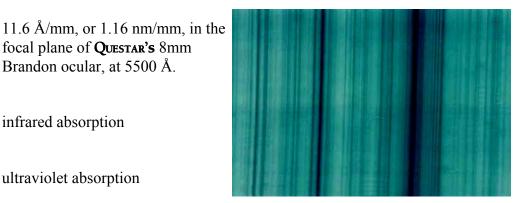
ultraviolet absorption

infrared absorption



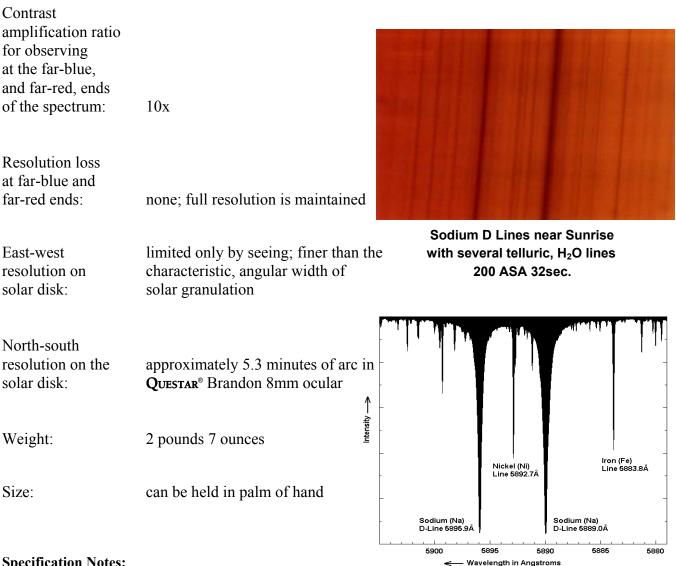


G-Band of CH (methyne)molecule at 4310Å, fully resolved. Color through eyepiece is rich royal blue.



H_γ at 4340Å with CH band in wing 31 Angstroms wide. 150 sec.

Film-photo spectrograms taken through 100mm lens at the QMAX[™] eyepiece Developed and printed at local 1-hr Photo shop. JPEG scans of prints.



- **Specification Notes:**
- 10 Angstrom units $(10\text{\AA}) = 1$ nanometer $(1 \text{ nm}) = 10^{-9}$ meters. **QUESTAR's** Maximum-Resolution 1. Solar Spectrometer just resolves the two faint lines at 4991.072 Å (Titanium) and 4991.275 Å (Iron), separated by 0.203 Å.
- 2. The optical relay preceding the ocular has a nominal 1:3 ratio. Thus, the ocular receives an image of the spectrum 3 times larger than the image formed by the spectrometer's imaging mirror. An 8mm ocular used as a magnifier results in a magnification of about 31.75, assuming 10" to be the closest a normal eve can approach an object. $3 \times 31.75 = 95.25$.
- The focal length of a 3.5" **OUESTAR**[®] telescope focused on the spectrometer's entrance slit is 3. approximately 1390mm. Since the spectrometer's collimating and imaging mirrors have the same focal length, no enlargement of the entrance slit occurs within the spectrometer. The tiny slice of the sun isolated by the entrance slit is delivered 1 to 1 into the image-plane of the spectrum, for a given wavelength. The 1:3 optical relay preceding the ocular then enlarges this tiny slice three times, including any portion of a sunspot that happens to lie focused on the entrance slit, exactly as if the focal length of a 3.5" QUESTAR[®] were 4170mm. An 8mm eyepiece used with a 4170mm focal length objective yields a power of $521.25 = 4170 \div 8$